

### **Developments in Group Field Theory Cosmology**

A collective effort: D. Oriti, E. Wilson-Ewing, S. Gielen, M. Sakellariadou, A. Pithis, M. de Cesare, A. Polaczek, A. Jercher, A. Calcinari, R. Dekhil, X. Pang, L. Mickel, T. Ladstätter, P. Fischer, ...

Luca Marchetti ILQGS 18 April 2023

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### Overview

### • Introduction to Group Field Theory

- · Group Field Theory and spinfoam models
- Group Field Theory and Loop Quantum Gravity
- Including scalar matter

### • Group Field Theory Cosmology

- Basic principles
- Homogeneous and isotropic sector
  - Volume
  - Matter
- Inhomogeneous sector
  - First steps and limitations
  - Super-horizon limit
  - Perturbations at all scales

### Introduction to Group Field Theory

Definition

Group Field Theories: theories of a field  $\varphi$  :  $G^d \to \mathbb{C}$  defined on *d* copies of a group manifold *G*. *d* is the dimension of the "spacetime to be" (*d* = 4) and *G* is the local gauge group of gravity,  $G = SL(2, \mathbb{C})$  or, in some cases, G = SU(2).

Oriti 1110.5606; Reisenberger, Rovelli 0002083; De Pietri, Petronio 0004045; Gurau 1006.0714; Baratin, Oriti 1002.4723; Finocchiaro, Oriti 1812.03550.

Definition

Action

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$$\mathcal{S}[arphi,ar{arphi}] = \int \mathrm{d}g_{s}ar{arphi}(g_{s})\mathcal{K}[arphi](g_{s}) + \sum_{\gamma}rac{\lambda_{\gamma}}{n_{\gamma}} \, \mathsf{Tr}_{\mathcal{V}\gamma}[arphi] + \mathsf{c.c.} \; .$$

- Interaction terms are combinatorially non-local.
- Field arguments convoluted pairwise following the combinatorial pattern dictated by the graph γ:

$$\mathsf{Tr}_{\mathcal{V}_{\gamma}}[\varphi] = \int \prod_{i=1}^{n_{\gamma}} \mathrm{d}g_{a} \prod_{(a,i;b,j)} \mathcal{V}_{\gamma}(g_{a}^{(i)}, g_{b}^{(j)}) \prod_{i=1}^{n_{\gamma}} \varphi(g_{a}^{(i)})$$



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$$Z[arphi,ar{arphi}] = \sum_{\Gamma} w_{\Gamma}(\{\lambda_{\gamma}\})A_{\Gamma}$$

- $\Gamma$  = stranded diagrams dual to *d*-dimensional cellular complexes of arbitrary topology.
- Amplitudes  $A_{\Gamma}$  = sums over group theoretic data associated to the cellular complex.

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Action

Partition function

Group Field Theories: theories of a field  $\varphi : G^d \to \mathbb{C}$  defined on d copies of a group manifold G. *d* is the dimension of the "spacetime to be" (d = 4) and *G* is the local gauge group of gravity,  $G = SL(2, \mathbb{C})$  or, in some cases, G = SU(2).

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$$Z[\varphi, \bar{\varphi}] = \sum_{\Gamma} w_{\Gamma}(\{\lambda_{\gamma}\})A_{\Gamma} = \text{ complete spinfoam model.}$$

- $\Gamma$  = stranded diagrams dual to *d*-dimensional cellular complexes of arbitrary topology.
- Amplitudes  $A_{\Gamma}$  = sums over group theoretic data associated to the cellular complex.
- $\blacktriangleright$   $\mathcal{K}$  and  $\mathcal{V}_{\gamma}$  chosen to match the desired spinfoam model.

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Action

The one-particle Hilbert space is  $\mathcal{H}_{tetra} \subset \otimes_{a=1}^4 \mathcal{H}_{\Delta_a}$  (subset defined by the imposition of constraints)

Finocchiaro, Oriti 1812.03550; Baez, Barrett 9903060; Baratin, Oriti 1002.4723; Gielen, Oriti 1004.5371; Oriti 1310.7786.

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Lie algebra (metric)
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$$\mathcal{H}_{\Delta_a} = L^2(\mathfrak{g})$$

Constraints

Geometricity constraints (appropriately encoded in  $\mathcal{K}$  and  $\mathcal{V}_{\gamma}$ ) allow for a d-1-simplicial interpretation of the fundamental quanta:

#### Closure

#### Simplicity

 $\sum_{a} B_{a} = 0$  (faces of the tetrahedron close).

•  $X \cdot (B - \gamma \star B)_a = 0$  (EPRL);

$$\blacktriangleright X \cdot B_a = 0 (BC).$$



Finocchiaro, Oriti 1812.03550; Baez, Barrett 9903060; Baratin, Oriti 1002.4723; Gielen, Oriti 1004.5371; Oriti 1310.7786.

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Lie algebra (metric)  $\mathcal{H}_{\Delta_a} = L^2(\mathfrak{g}) \xleftarrow{\text{Non-comm.}}{\text{FT}} \qquad \mathcal{H}_{\Delta_a} = L^2(G)$ Constraints

Geometricity constraints (appropriately encoded in  $\mathcal{K}$  and  $\mathcal{V}_{\gamma}$ ) allow for a d-1-simplicial interpretation of the fundamental quanta:

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Developments in GFT Cosmology

g4

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d - 1-simplicial interpretation of the fundamental quanta:

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ClosureSimplicity $j_3$  $\sum_a B_a = 0$  $\blacktriangleright X \cdot (B - \gamma \star B)_a = 0$  (EPRL);(faces of the tetrahedron close). $\blacktriangleright X \cdot B_a = 0$  (BC).



- Impose simplicity and reduce to G = SU(2).
- Impose closure (gauge invariance).

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LQG

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Geometricity constraints (appropriately encoded in  $\mathcal{K}$  and  $\mathcal{V}_{\gamma}$ ) allow for a d-1-simplicial interpretation of the fundamental quanta:

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$$j_3$$
  $j_1$   $j_1$ 

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  - Impose closure (gauge invariance).

$$\begin{aligned} \mathcal{H}_{\mathsf{tetra}} &= \bigoplus_{\vec{j}} \mathsf{Inv} \left[ \bigotimes_{a=1}^{4} \mathcal{H}_{j_a} \right] \\ &= \mathsf{open \ spin-network \ vertex \ space} \end{aligned}$$

Finocchiaro, Oriti 1812.03550: Baez, Barrett 9903060: Baratin, Oriti 1002.4723: Gielen, Oriti 1004.5371: Oriti 1310.7786.

Developments in GFT Cosmology

Simplicity

LQG

#### Tetrahedron wavefunction

 $\varphi(g_1,\ldots,g_4)$  (subject to constraints)

Oriti 1310.7786; Oriti 1408.7112; Sahlman, Sherif 2302.03612.



Oriti 1310.7786; Oriti 1408.7112; Sahlman, Sherif 2302.03612.



$$\mathcal{F}_{\mathsf{GFT}} = \bigoplus_{V=0}^{\infty} \operatorname{sym} \left[ \mathcal{H}_{\mathsf{tetra}}^{(1)} \otimes \mathcal{H}_{\mathsf{tetra}}^{(2)} \otimes \ldots \mathcal{H}_{\mathsf{tetra}}^{(V)} \right]$$

- ▶  $\mathcal{F}_{GFT}$  generated by action of  $\hat{\varphi}^{\dagger}(g_a)$  on  $|0\rangle$ , with  $[\hat{\varphi}(g_a), \hat{\varphi}^{\dagger}(g'_a)] = \mathbb{I}_G(g_a, g'_a)$ .
- $\mathcal{H}_{\Gamma} \subset \mathcal{F}_{GFT}$ ,  $\mathcal{H}_{\Gamma}$  space of states associated to connected simplicial complexes  $\Gamma$ .
- Generic states do not correspond to connected simplicial lattices nor classical simplicial geometries.
- ▶ Similar to *H*<sub>LQG</sub> but also different: no continuum intuition, orthogonality wrt nodes, not graphs.

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Volume operator 
$$\hat{V} = \int dg_a^{(1)} dg_a^{(2)} V(g_a^{(1)}, g_a^{(2)}) \hat{\varphi}^{\dagger}(g_a^{(1)}) \hat{\varphi}(g_a^{(2)}) = \sum_{j_a, m_a, \iota} V_{j_a, \iota} \hat{\varphi}^{\dagger}_{j_a, m_a, \iota} \hat{\varphi}_{j_a, m_a, \iota}$$

Generic second quantization prescription to build a m + n-body operator: sandwich matrix elements between spin-network states between m powers of φ<sup>2</sup> and n powers of φ<sup>2</sup>.

Oriti 1310.7786; Oriti 1408.7112; Sahlman, Sherif 2302.03612.

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Operators

Group Field Theories: theories of a field  $\varphi: G^d \to \mathbb{C}$  defined on the product  $G^d$ .

 $\begin{aligned} d \text{ is the dimension of the "spacetime to be"} & (d = 4) \\ & \text{and } G \text{ is the local gauge group of gravity,} \\ & G = \mathrm{SL}(2,\mathbb{C}) \text{ or, in some cases, } G = \mathrm{SU}(2). \end{aligned}$ 

#### Kinematics

Quanta are d-1-simplices decorated with quantum geometric and scalar data:

Geometricity constraints imposed analogously as before.

### Dynamics

 $S_{GFT}$  obtained by comparing  $Z_{GFT}$  with simplicial gravity + scalar fields path integral.

 Geometric data enter the action in a non-local and combinatorial fashion.



Li, Oriti, Zhang 1701.08719; Oriti 0912.2441; Gielen, Oriti 1311.1238; Oriti, Sindoni, Wilson-Ewing 1602.05881; Gielen, Sindoni 1602.08104; ...

Group Field Theories: theories of a field  $\varphi$  :  $G^d \times \mathbb{R}^{d_l} \to \mathbb{C}$  defined on the product of  $G^d$  and  $\mathbb{R}^{d_l}$ . d is the dimension of the "spacetime to be" (d = 4)and G is the local gauge group of gravity,  $G = SL(2, \mathbb{C})$  or, in some cases, G = SU(2).

#### Kinematics

Quanta are d - 1-simplices decorated with quantum geometric and scalar data:

- Geometricity constraints imposed analogously as before.
- Scalar field discretized on each *d*-simplex: each *d* − 1-simplex composing it carries values *x* ∈ ℝ<sup>d</sup><sub>1</sub>.

### Dynamics

 $S_{GFT}$  obtained by comparing  $Z_{GFT}$  with simplicial gravity + scalar fields path integral.

- Geometric data enter the action in a non-local and combinatorial fashion.
- Scalar field data are local in interactions.
- ▶ For minimally coupled, free, massless scalars:

 $\mathcal{K}(g_a, g_b; \chi^{\alpha}, \chi^{\alpha'}) = \mathcal{K}(g_a, g_b; (\chi^{\alpha} - \chi^{\alpha'})^2)$  $\mathcal{V}_5(g_a^{(1)}, \dots, g_a^{(5)}, \chi) = \mathcal{V}_5(g_a^{(1)}, \dots, g_a^{(5)})$ 

Li, Oriti, Zhang 1701.08719; Oriti 0912.2441; Gielen, Oriti 1311.1238; Oriti, Sindoni, Wilson-Ewing 1602.05881; Gielen, Sindoni 1602.08104; ...



### Group Field Theory Cosmology

**Collective states** 

### **GFT** condensates

From the GFT perspective, continuum geometries are associated to large number of quanta.

The simplest states that can accommodate infinite number of quanta are condensate states:

$$|\sigma\rangle = \mathcal{N}_{\sigma} \exp\left[\int \mathrm{d}^{d_{f}} \chi \int \mathrm{d}g_{\mathfrak{s}} \,\sigma(g_{\mathfrak{s}},\chi^{\alpha}) \hat{\varphi}^{\dagger}(g_{\mathfrak{s}},\chi^{\alpha})\right] |0\rangle$$

LM, Oriti 2008.02774; LM, Oriti 2112.12677; Oriti, Sindoni, Wilson-Ewing 1602.05881; Gielen, Oriti, Sindoni 1311.1238; Gielen 1404.2944

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### Mean-field approximation

• When interactions are small (certainly satisfied in an appropriate regime) the dynamics of  $\sigma$  is:

$$\left\langle \frac{\delta S[\hat{\varphi}, \hat{\varphi}^{\dagger}]}{\delta \hat{\varphi}(g_{I}, x^{\alpha})} \right\rangle_{\sigma} = \int \mathrm{d}h_{a} \int \mathrm{d}\chi \, \mathcal{K}(g_{a}, h_{a}, (x^{\alpha} - \chi^{\alpha})^{2}) \sigma(h_{a}, \chi^{\alpha}) + \lambda \frac{\delta V[\varphi, \varphi^{*}]}{\delta \varphi^{*}(g_{a}, x^{\alpha})} \bigg|_{\varphi = \sigma} = 0 \,.$$

▶ Non-perturbative: equivalent to a mean-field (saddle-point) approximation of Z.

LM, Oriti 2008.02774; LM, Oriti 2112.12677; Oriti, Sindoni, Wilson-Ewing 1602.05881; Gielen, Oriti, Sindoni 1311.1238; Gielen 1404.2944

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### **Condensate Peaked States**

Constructing relational observables on *F*<sub>GFT</sub> is difficult (QFT with no continuum intuition).

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**Collective states** 

Relationality

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#### **Condensate Peaked States**

- Constructing relational observables on  $\mathcal{F}_{GFT}$  is difficult (QFT with no continuum intuition).
- Relational localization implemented at an effective level on observable averages.
- If  $\chi^{\mu}$  constitute a reference frame, this can be achieved by assuming

 $\sigma = (\text{fixed peaking function } \eta) \times (\text{dynamically determined reduced wavefunction } \tilde{\sigma})$ 

LM, Oriti 2008.02774; LM, Oriti 2112.12677; Oriti, Sindoni, Wilson-Ewing 1602.05881; Gielen, Oriti, Sindoni 1311.1238; Gielen 1404.2944.

Homogeneous sector

Spatial relational homogeneity:

 $\sigma$  depends on a MCMF "clock" scalar field  $\chi^0$ (D = minisuperspace + clock)

Oriti 2008.02774; LM, Oriti 2010.09700; Oriti, Sindoni, Wilson-Ewing 1602.05881; Jercher, Oriti, Pithis 2112.00091.

Spatial relational homogeneity:  $\sigma$  depends on a MCMF "clock" scalar field  $\chi^0$ (D = minisuperspace + clock)

### **Collective Observables**

Number, volume (determined e.g. by the mapping with LQG) and matter operators (notation:  $(\cdot, \cdot) = \int d\chi^0 dg_a$ ):

$$\begin{split} \hat{N} &= (\hat{\varphi}^{\dagger}, \hat{\varphi}) & \hat{V} &= (\hat{\varphi}^{\dagger}, V[\hat{\varphi}]) \\ \hat{\chi}^{0} &= \left(\hat{\varphi}^{\dagger}, \chi^{0} \hat{\varphi}\right) & \hat{\Pi}^{0} &= -i(\hat{\varphi}^{\dagger}, \partial_{0} \hat{\varphi}) \end{split}$$

• Observables  $\leftrightarrow$  collective operators on Fock space.

Oriti 2008.02774; LM, Oriti 2010.09700; Oriti, Sindoni, Wilson-Ewing 1602.05881; Jercher, Oriti, Pithis 2112.00091.

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- $\begin{array}{l} \blacktriangleright \quad \langle \hat{O} \rangle_{\sigma_{\chi^0}} = O[\tilde{\sigma}]|_{\chi^0 = x^0}: \\ \text{functionals of } \tilde{\sigma} \\ \text{localized at } x^0. \end{array}$

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Int

### **Collective Observables**

Number, volume (determined e.g. by the mapping with LQG) and matter operators (notation:  $(\cdot, \cdot) = \int d\chi^0 dg_a$ ):

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- ▶ Observables ↔ collective operators on Fock space.
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Developments in GFT Cosmology

#### Relationality

Averaged evolution wrt x<sup>0</sup> is physical:

- Emergent effective relational description:
  - Small clock quantum fluctuations.
  - Effective Hamiltonian  $H_{\sigma_{\chi^0}} \simeq \langle \hat{\Pi}^0 \rangle_{\sigma_{\chi^0}}$ .

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▶ Observables ↔ collective operators on Fock space.

#### Relationality

Averaged evolution wrt x<sup>0</sup> is physical:

$$\langle \hat{\chi}^0 \rangle_{\sigma_{\chi^0}} \equiv \langle \hat{X}^0 \rangle_{\sigma_{\chi^0}} / \langle \hat{N} \rangle_{\sigma_{\chi^0}} \simeq x^0$$

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  - Small clock quantum fluctuations.
  - Effective Hamiltonian  $H_{\sigma_{\chi^0}} \simeq \langle \hat{\Pi}^0 \rangle_{\sigma_{\chi^0}}$ .

$$\begin{array}{l} \langle \hat{O} \rangle_{\sigma_{\chi^{0}}} = O[\tilde{\sigma}]|_{\chi^{0} = x^{0}}: \\ \text{functionals of } \tilde{\sigma} \\ \text{localized at } x^{0}. \end{array} \begin{array}{l} \text{Wavefunction} \\ \text{isotropy} \end{array} \qquad \begin{array}{l} \langle \hat{V} \rangle_{\sigma_{\chi}^{0}} = \sum_{\upsilon}^{f} |\nabla_{\upsilon}|^{2} \langle x^{0} \rangle \\ \langle \hat{N} \rangle_{\sigma_{\chi}^{0}} = \sum_{\upsilon}^{f} |\sigma_{\upsilon}|^{2} \langle x^{0} \rangle \end{array} \begin{array}{l} \nu = j \in \mathbb{N}/2 \text{ (EPRL)}; \\ \nu = \rho \in \mathbb{R} \text{ (ext. BC)}. \end{array}$$

Oriti 2008.02774; LM, Oriti 2010.09700; Oriti, Sindoni, Wilson-Ewing 1602.05881; Jercher, Oriti, Pithis 2112.00091.



#### Mean-field approximation

- ▶ Mesoscopic regime: large *N* but negligible interactions.
- Derivative expansion of  $\mathcal{K}$  (due to peaking properties).
- Isotropy:  $\tilde{\sigma}_{\upsilon} \equiv \rho_{\upsilon} e^{i\theta_{\upsilon}}$  fundamental variables.

 $\tilde{\sigma}_{\upsilon}^{\prime\prime}-2i\tilde{\pi}_{0}\tilde{\sigma}_{\upsilon}^{\prime}-E_{\upsilon}^{2}\tilde{\sigma}=0.$ 

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 $\frac{\text{Effective relational Freidmann dynamics}}{\left(\frac{V'}{3V}\right)^2 \simeq \left(\frac{2 \, \text{\pounds}_{\upsilon} \, V_{\upsilon} \rho_{\upsilon} \operatorname{sgn}(\rho'_{\upsilon}) \sqrt{\mathcal{E}_{\upsilon} - \mathcal{Q}_{\upsilon}^2 / \rho_{\upsilon}^2 + \mu_{\upsilon}^2 \rho_{\upsilon}^2}}{3 \, \text{\pounds}_{\upsilon} \, V_{\upsilon} \rho_{\upsilon}^2}\right)^2, \quad \frac{V''}{V} \simeq \frac{2 \, \text{\pounds}_{\upsilon} \, V_{\upsilon} \left[\mathcal{E}_{\upsilon} + 2\mu_{\upsilon}^2 \rho_{\upsilon}^2\right]}{\text{\pounds}_{\upsilon} \, V_{\upsilon} \rho_{\upsilon}^2}$ 

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#### Classical limit (large $\rho_v$ s, late times)

If μ<sup>2</sup><sub>v</sub> is mildly dependent on v (or one v is dominating) and equal to 3πG

 $(V'/3V)^2 \simeq 4\pi G/3 \longrightarrow \text{flat FLRW}$ 

 Quantum fluctuations on clock and geometric variables are under control.

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 Quantum fluctuations on clock and geometric variables are under control.

#### Bounce

- A non-zero volume bounce happens for a large range of initial conditions (at least one Q<sub>v</sub> ≠ 0 or one E<sub>v</sub> < 0).</p>
- The average singularity resolution may still be spoiled by quantum effects on geometric and clock variables.

LM, Oriti 2008.02774; LM, Oriti 2010.09700; Oriti, Sindoni, Wilson-Ewing 1602.05881; Jercher, Oriti, Pithis 2112.00091; ...

### (T)GFT interactions and matter

#### Running couplings and effective potentials

- Adding a scalar field φ with potential U<sub>φ</sub> requires (T)GFT interactions, as V<sub>γ</sub> = V<sub>γ</sub>({g}, U<sub>φ</sub>).
- Interactions studied perturbatively at late times (mesoscopic regime) and in single j approx.

Ladstätter, LM, Oriti (to appear); Oriti, Pang 2105.03751.
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**Modulus interactions** 

notation: 
$$(\cdot, \cdot) = \int d^4 \chi d\phi dg_a$$
  
 $\operatorname{Fr}_{\mathcal{V}_{\gamma_l}}^{(m)} [\varphi, \bar{\varphi}] \sim (\mathcal{V}_{\gamma_l}^{(m)}, \bar{\varphi}^{(l+1)/2} \varphi^{(l+1)/2})$ 

 $\checkmark$  GR matching possible only if l = 5, and if

Macroscopic constants (including G) run with relational time!

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### Phase interactions

$$\begin{split} \text{notation:} & (\cdot, \cdot) = \int \mathrm{d}^4 \chi \mathrm{d} \phi \mathrm{d} g_{\text{a}} \\ & \mathsf{Tr}_{\mathcal{V}\gamma_l}^{(p)} = \big( \mathcal{V}_{\gamma_l}^{(p)}, \varphi^{l+1} \big) \end{split}$$

- $\checkmark$  GR matching possible only if l = 5, but
- Effective scalar field potential corrected by trigonometric factors.

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### Phase interactions

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# Phantom dark energy

### **Emergent matter components**

- Matter can also emerge as a result of pure QG effects!
- Consider modulus interactions at very late times, but include a subdominant spin j':

$$w = 3 - 2(VV'')/(V')^2 \simeq -1 - b/V$$
,  $b > 0$ .

· Universe effectively dominated by (non-pathologic) emergent phantom dark energy.

Ladstätter, LM, Oriti (to appear); Oriti, Pang 2105.03751.

Luca Marchetti

Inhomogeneous sector



Gielen, Oriti, 1709.01095; Gerhart, Oriti, Wilson-Ewing 1805.03099.



► Fock quantize the emergent perturbation dynamics.

Gielen, Oriti, 1709.01095; Gerhart, Oriti, Wilson-Ewing 1805.03099.



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  - First step in GFT using second moment  $\delta^2 \hat{V}$ .
  - What about higher moments? Is V really relational? Only a background result?

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Developments in GFT Cosmology

► Goal: vector, tensor, scalar modes at all scales.



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  - Can we extend to all scales? Can we use a proper physical reference frame?

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Simplest (slightly) relationally inhomogeneous system

# Classical

- ► 4 MCMF reference fields (\(\chi^0\), \(\chi^i\)), with Lorentz/Euclidean invariant \(S\_\chi\$ in field space.
- 1 MCMF matter field φ dominating the e.m. budget and relationally inhomog. wrt. χ<sup>i</sup>.

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## Quantum

- ► GFT field: φ(g<sub>a</sub>, χ<sup>μ</sup>, φ), depends on 5 discretized scalar variables.
- EPRL-like model with S<sub>GFT</sub> respecting the classical matter symmetries.

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# Observables

# Aat. Vol. Frame

$$\begin{array}{l} \text{notation:} (\cdot, \cdot) = \int \mathrm{d}^4 \chi \mathrm{d}\phi \mathrm{d}g_a \\ \\ \hat{X}^{\mu} = (\hat{\varphi}^{\dagger}, \chi^{\mu} \hat{\varphi}) \quad \hat{\Pi}^{\mu} = -i(\hat{\varphi}^{\dagger}, \partial_{\mu} \hat{\varphi}) \\ \\ \text{Only isotropic info:} \quad \hat{V} = (\hat{\varphi}^{\dagger}, V[\hat{\varphi}]) \end{array}$$

 $\hat{\Phi} = (\hat{\varphi}^{\dagger}, \phi \hat{\varphi}) \qquad \hat{\Pi}_{\phi} = -i(\hat{\varphi}^{\dagger}, \partial_{\phi} \hat{\varphi})$ 

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### States

- CPSs around  $\chi^{\mu} = x^{\mu}$ , with
  - η: Isotropic peaking on rods;
  - *σ*: Isotropic distribution of geometric data.
- Small relational  $\tilde{\sigma}$ -inhomogeneities ( $\tilde{\sigma} = \rho e^{i\theta}$ ):

 $\rho = \bar{\rho}(\cdot, \chi^{0}) + \delta\rho(\cdot, \chi^{\mu}), \ \theta = \bar{\theta}(\cdot, \chi^{0}) + \delta\theta(\cdot, \chi^{\mu}).$ 

LM, Oriti 2112.12677; Gerhart, Oriti, Wilson-Ewing 1805.03099.

Luca Marchetti

Aat. Vol. Frame









# Super-horizon scalar perturbations



Bertschinger 0604485; Fischer, LM, Oriti (to appear); LM, Oriti 2112.12677; Gielen, Mickel 2211.04500.

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	Super-horizon volume and matter dynamics
<ul> <li>Averaged q.e.o.m. (no interactions) → co</li> <li>Restrict to super-horizon modes but study a</li> </ul>	bupled eqs. for $(\rho, \theta)$ . also early times. $\begin{array}{c} \text{single} \\ \text{spin} \end{array}  \begin{array}{c} \text{Dynamic equations} \\ \text{for } \langle \hat{V} \rangle_{\sigma}, \langle \hat{\Phi} \rangle_{\sigma} \end{array}$
Modified gravity	Porturbing background dynamics

### Modified gravity

- Dynamics of super-horizon scalar perturbations can be obtained generically for any MG theory.
- No matching at early times with effective GFT volume dynamics.

## Perturbing background dynamics

- Study super-horizon scalar perturbations by perturbing background QG volume eq.
- No matching at early times with full effective GFT volume dynamics

Bertschinger 0604485; Fischer, LM, Oriti (to appear); LM, Oriti 2112.12677; Gielen, Mickel 2211.04500.

## Luca Marchetti

# Scalar perturbations at all scales

# Causal frame fields coupling

Causal properties of frame fields can be easily implemented in the complete extended BC model.

•  $\varphi_{\alpha} \equiv \varphi(g_a, X_{\alpha}, \chi^{\mu}, \phi), g \in SL(2, \mathbb{C}), X_{\alpha}$  tetrahedron normal defining its causal character,  $\alpha = \pm$ .

Jercher, LM, Pithis (to appear); Jercher, Oriti, Pithis 2206.15442.

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**Two-sector Fock space** 

- Generic operators on  $\mathcal{F} = \mathcal{F}_+ \otimes \mathcal{F}_-$  correlate spacelike and timelike tetrahedra.
- ▶ Volume operator is an exception:  $\hat{V} = \hat{V}_+ \otimes \mathbb{I}_-$ .

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$$\begin{split} \mathcal{K}_{+} &= \mathcal{K}_{+}(\cdot, (\chi^{0}-\chi^{0\prime})^{2})\,,\\ \mathcal{K}_{-} &= \mathcal{K}_{-}(\cdot, |\boldsymbol{\chi}-\boldsymbol{\chi}'|^{2})\,. \end{split}$$

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**Collective states** 

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Including two-body correlations

$$|\psi
angle = \mathcal{N}_{\psi} \exp(\hat{\sigma} \otimes \mathbb{I}_{-} + \mathbb{I}_{+} \otimes \hat{\tau} + \widehat{\delta \Phi} \otimes \mathbb{I}_{-} + \widehat{\delta \Psi} + \mathbb{I}_{+} \otimes \widehat{\delta \Xi}) |0
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### Background

- $\hat{\sigma} = (\sigma, \hat{\varphi}^{\dagger}_{+})$ : spacelike condensate.
- $\hat{\tau} = (\tau, \hat{\varphi}_{-}^{\dagger})$ : timelike condensate.
- τ, σ peaked; τ̃, σ̃ homogeneous.

Jercher, LM, Pithis (to appear); Jercher, Oriti, Pithis 2206.15442.

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### Background

• 
$$\hat{\sigma} = (\sigma, \hat{\varphi}_+^{\dagger})$$
: spacelike condensate.

- $\hat{\tau} = (\tau, \hat{\varphi}^{\dagger}_{-})$ : timelike condensate.
- $\tau$ ,  $\sigma$  peaked;  $\tilde{\tau}$ ,  $\tilde{\sigma}$  homogeneous.

## Perturbations

- $\bullet \quad \widehat{\delta\Phi} = (\delta\Phi, \hat{\varphi}_{+}^{\dagger}\hat{\varphi}_{+}^{\dagger}), \ \widehat{\delta\Psi} = (\delta\Psi, \hat{\varphi}_{+}^{\dagger}\hat{\varphi}_{-}^{\dagger}), \ \widehat{\delta\Xi} = (\delta\Xi, \hat{\varphi}_{-}^{\dagger}\hat{\varphi}_{-}^{\dagger}).$
- $\delta \Phi$ ,  $\delta \Psi$  and  $\delta \Xi$  small and relationally inhomogeneous.
- Perturbations = nearest neighbour 2-body correlations.

Jercher, LM, Pithis (to appear); Jercher, Oriti, Pithis 2206.15442.

Luca Marchetti



# Scalar perturbations

Mean-field equations (negligible interactions):

$$\left<\delta S/\delta\hat{\varphi}_{+}^{\dagger}\right>_{\psi} = 0 = \left<\delta S/\delta\hat{\varphi}_{-}^{\dagger}\right>_{\psi}$$

- 2 coupled eqs. for 3 variables: (δΦ, δΨ, δΞ)!
- Late times and single (spacelike) rep. label.

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- Late time GR matching fixes:
  - Parameters determining τ dynamics;
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  - Dynamical freedom (e.g. in  $\delta \Phi$ ).

Late times volume perturbations dynamics matches GR at all scales!

Jercher, LM, Pithis (to appear); Jercher, Oriti, Pithis 2206.15442.





- Singularity resolution into quantum bounce.
- Universal bounce (for MCMF scalar field).
- Impact of quantum effects on the bounce (and interplay with relationality).
- Acceleration produced by the bounce not long enough to sustain inflation.

LM, Oriti 2008.02774 - 2010.09700; Oriti, Sindoni, Wilson-Ewing 1602.05881; Jercher, Oriti, Pithis 2112.00091; de Cesare, Pithis, Sakellariadou 1606.00352; Ladstätter, LM, Oriti (to appear); Oriti, Pang 2105.03751; Gielen, Polaczek 1912.06143 ; ...

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- Singularity resolution into quantum bounce.
- Universal bounce (for MCMF scalar field).
- Impact of quantum effects on the bounce (and interplay with relationality).
- Acceleration produced by the bounce not long enough to sustain inflation.

- Small interactions: classical regime identified (small quantum fluctuations and GR matching).
- ✓ Universal classical limit (for MCMF scalar field).
- Inclusion of scalar field with potential: emergent running couplings.
- Exotic matter can emerge from GFT interactions.

LM, Oriti 2008.02774 - 2010.09700; Oriti, Sindoni, Wilson-Ewing 1602.05881; Jercher, Oriti, Pithis 2112.00091; de Cesare, Pithis, Sakellariadou 1606.00352; Ladstätter, LM, Oriti (to appear); Oriti, Pang 2105.03751; Gielen, Polaczek 1912.06143 ; ...

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- Extend the analysis to more generic fluids.
- Universal bounce also for generic fluids?
- In particular, would trigonometric modifications to a scalar field potential appear at early times?
- Mhat kind of inflationary physics is generated?

- Extend the analysis to more generic fluids.
- Universal classical limit also for generic fluids?
- Insights on the renormalization properties of GFTs from emergent running couplings?
- Can we rely on mean-field approx. at late times?




### Results

- Super-horizon analysis in EPRL with MCMF scalar fields:
  - Scalar pert. dynamics differs from any MG model.
  - Full QG scalar pert. dynamics differs from perturbed background dynamics.

LM, Oriti 2112.12677; Fischer, LM, Oriti (to appear); Jercher, LM, Pithis (to appear); Gerhart, Oriti, Wilson-Ewing 1805.03099.



- Full QG scalar pert. dynamics differs from perturbed background dynamics.
- $\checkmark$  Scalar pert.  $\longleftrightarrow$  guantum correlations!
- ✓ Late-times scalar pert. dynamics matches GR!

LM, Oriti 2112.12677; Fischer, LM, Oriti (to appear); Jercher, LM, Pithis (to appear); Gerhart, Oriti, Wilson-Ewing 1805.03099.

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Developments in GFT Cosmology



#### Perspectives

- ▲ Different fundamental d.o.f. → different perturbation dynamics?
- ▲ Scalar field perturbations? EFT description?
- Are the results universal? Analysis in BC!
- Generalization to physically interesting fluids.
- Extension to VT modes: more observables!
- Initial conditions and power spectra?
  - Fock quantization of early-times dynamics.
  - Can we derive it from full QG?

### Perspectives

- Physical interpretation and consequences of matching conditions?
- Scalar field perturbations? EFT description?
- Are the results universal? Extension to EPRL!
- Generalization to physically interesting fluids.
- Extension to VT modes: more observables!
- How do quantum perturbations classicalize?
- How do GFT interactions change the picture?

Fischer, LM, Oriti (to appear); Jercher, LM, Pithis (to appear); Dekhil, Liberati, Oriti (to appear); Calcinari, Gielen 2210.03149.

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# Backup

## **Specifics of GFT models**

$$\begin{split} S &= \sum_{\{j_{a}\}, \{j_{a}'\}, \{m_{a}\}, \{m_{a}'\}, \{n_{a}'\}, (\mu_{a}'), \iota, \iota'} \bar{\varphi}_{\{m_{a}\}}^{\{j_{a}\}, \iota'} \mathcal{K}_{\{m_{a}\}}^{\{j_{a}\}, \{j_{a}'\}, \iota'} + V_{5}, \\ V_{5} &= \frac{1}{5} \sum_{\{j_{a}\}, \{m_{a}\}, \{\iota_{b}\}} \varphi_{m_{1}m_{2}m_{3}m_{4}}^{(j_{1}j_{3})j_{1}j_{1}} \mathcal{K}_{-m_{4}m_{5}m_{7}}^{(j_{1}j_{3})j_{1}j_{3}} \mathcal{K}_{-m_{7}-m_{3}m_{8}m_{9}}^{(j_{2}j_{5})j_{1}0\iota_{4}} \mathcal{K}_{-m_{1}0-m_{8}-m_{5}-m_{1}} \\ &\times \prod_{c=1}^{10} (-1)^{j_{c}-m_{c}} \mathcal{V}_{5}(j_{1}, \ldots, j_{10}; \iota_{1}, \ldots, \iota_{5}), & a = 1, \ldots, 4 \\ &b = 1, \ldots, 5 \\ \mathcal{V}_{5}(\{j_{c}\}, \{\iota_{b}\}) &= \sum_{\{n_{A}\}} \int \left[ \prod_{A} d\rho_{A}(n_{A}^{2} + \rho_{A}^{2}) \right] \left[ \bigotimes_{b} f_{\{n_{A}\}\{\rho_{A}\}}^{\iota_{b}}(\{j_{a}\}) \right] \{15j\}_{\mathrm{SL}(2,\mathbb{C})}, \end{split}$$

where f maps  $SL(2, \mathbb{C})$  data into SU(2) ones by imposing the constraints n = 2j and  $\rho = 2j\gamma$ .

$$\begin{split} S &= \left[\prod_{i} \int d\rho_{i} \, 4\rho_{i}^{2} \sum_{j_{i}m_{i}}\right] \bar{\varphi}_{j_{i}m_{i}}^{\rho_{i}} \varphi_{j_{i}m_{i}}^{\rho_{i}} + \frac{\lambda}{5} \left[\prod_{a=1}^{10} \int d\rho_{a} \, 4\rho_{a}^{2} \sum_{j_{a}m_{a}}\right] \left[\prod_{a=1}^{10} (-1)^{-j_{a}-m_{a}}\right] \{10\rho\}_{BC} \right] \\ &\times \varphi_{j_{1}m_{1}j_{2}m_{2}j_{3}m_{3}m_{4}m_{4}}^{\rho_{1}\rho_{2}\rho_{5}\rho_{6}\rho_{7}} \varphi_{j_{7}-m_{7}j_{3}-m_{3}j_{8}m_{8}j_{9}m_{9}} \\ &\times \varphi_{j_{9}-m_{9}j_{6}-m_{6}j_{2}-m_{2}j_{10}m_{10}}^{\rho_{10}\rho_{10}\rho_{5}\rho_{11}} \varphi_{j_{10}-m_{10}j_{8}-m_{8}j_{5}-m_{5}j_{1}-m_{1}} + c.c. \end{split}$$

Engle, Livine, Pereira, Rovelli 0711.0146; Gielen, Oriti, Sindoni 1311.1238; Jercher, Oriti, Pithis 2112.00091.

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EPRL model

Extended BC model

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